

# Mars 2020 Perseverance Rover Status

Ken Farley, Caltech Mars 2020 Project Scientist

Katie Stack Morgan, JPL Mars 2020 Deputy Project Scientist

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**Jet Propulsion Laboratory**California Institute of Technology

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## **Mission Status**

- Planning sol 339 on Mars
- The rover is healthy and all science instruments except MEDA are functioning with no significant issues
- 18 Ingenuity helicopter flights
- Total mission odometry: 2914 m
- Abrasions: 4
- Sample Tubes Sealed: 8
  - 3 pairs of rock cores
  - 1 atmospheric sample
  - 1 witness tube



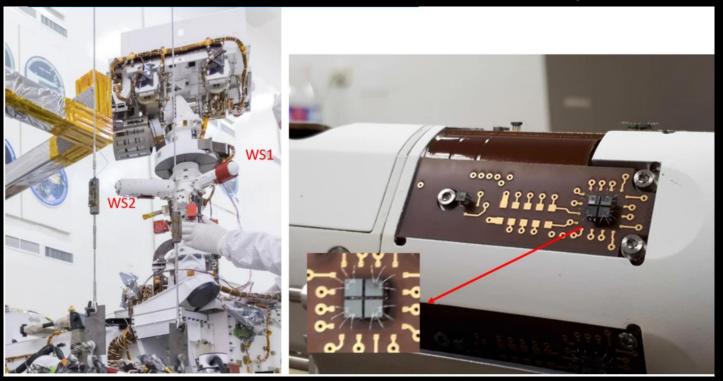
# Two PDS Deliveries Completed

| Release | Acquisition (Sols) | Release to Public |
|---------|--------------------|-------------------|
| 1       | 0 - 89             | 20-Aug-21         |
| 2       | 90 - 179           | 22-Nov-21         |
| 3       | 180 - 299          | 22-Mar-22         |
| 4       | 300 - 419          | 22-Jul-22         |
| 5       | 420 - 539          | 21-Nov-22         |
| 6       | 540-639            | 7-Mar-23          |
| 7       | 640 - 669          | 6-Jul-23          |

Nothing is ever easy (on Mars): Sol 306 "Pebble Anomaly"



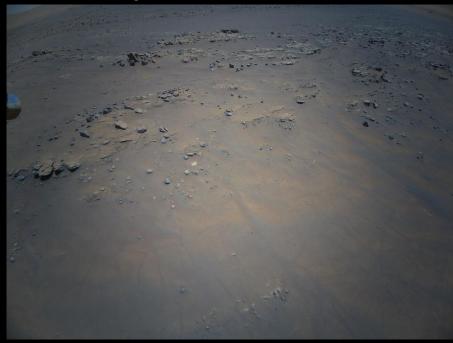
## MEDA Wind Sensor Anomaly

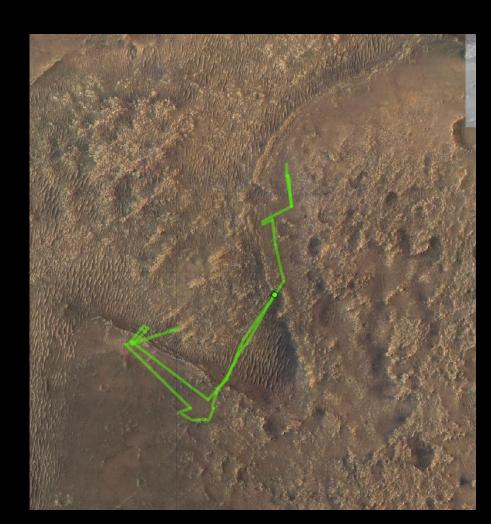


Several of the wind sensor transducers experienced faults
Assessment of the cause and implications of these faults is in progress

# **Ingenuity Flights**

- 18 flights as of February 2
- heli now being staged ahead of rover in preparation for drive campaign to delta





## MOXIE

- 8 Oxygen generation runs have been performed, all successfully
  - including worst-case low-pressure conditions
- O<sub>2</sub> purity is close to 100% when operated with that objective
- Degradation of the electrolytic stack is very low, many more runs are possible
- Upcoming runs will be more diagnostic in nature, refining understanding of the operational characteristics of the device

#### Sampling Details

#### How are Samples Selected?

Sample priority guided by community desires expressed in published documents, e.g., iMOST Report, as interpreted and applied to Jezero crater by the Mars 2020 Science Team

#### What "Field Notes" are being acquired?

Every sample (or paired sample set) is documented by a standardized set of observations that are executed following an optimized sol path of 10-15 sols duration. This Standardized Observation Protocol (STOP list) maximizes efficiency and consistency.

#### Included:

- 1) workspace imaging at multiple scales using (ZCAM, ECAM, SCAM RMI, WATSON)
- 2) workspace and abrasion-patch proximity and remote science (SHERLOC, PIXL, SCAM, ZCAM)
- 3) borehole and cuttings imaging and remote science (WATSON, SCAM, ZCAM)

Additional opportunistic science is also undertaken in association with sample collection.

#### Sampling Details

#### How are sample-related data recorded?

#### 1) Sample Dossier

A digital "one stop shop" file that includes

- 1) Links to all STOP list observations for a given sample stored in PDS
- 2) Additional rover-related data associated with sampling, for example
  - Rover localization
  - Sampling event time history and coring operational and sensor details
  - Temperatures
  - Core length estimate and Cachecam images

#### Sampling Details

#### How are sample-related data recorded?

#### 2) Initial Reports

A templatized narrative description of each sample, written by the science team within 3 weeks of collection

- Why the sample was collected
- How the sample fits in geologic context
- Description and initial interpretation of STOP list data
- Initial interpretation of the sample and its history
- Assessment of likely uses of the sample if returned to Earth

This document is preliminary, and is not revised after completion...

Detailed interpretation developed over time will appear in peer reviewed publications

## Example pages from Roubion Initial Report

M2020-164-2 Roubion

#### INITIAL REPORT

#### M2020-164-2 Roubion

(no core recovered, atmospheric sample)

Sample Designation: M2020-164-2 Roubion

Date of Coring: 5-Aug-2021

Mars Time of Sample Core Sealing: 19:11:35 LMST, Sol 164, Ls 81.1

Latitude (N), Longitude (E), Elevation: 18.42769340, 77.45165066, -2584.96 m

Campaign: Crater Floor

Region of Interest: Séitab Thumb

Lithology: Fine- to medium-grained mafic and likely igneous rock, possibly basalt or microgabbro (alternatively, basaltic sandstone). Primary minerals are plagioclase and pyroxene, also possibly apatite and EET, oxides. Weathering and/or aqueous alteration is indicated by pits and crevices in abraded surface and abundant secondary minerals including iron oxide (possibly hydrated), sulfates, perchlorate, and possibly phosphate and halite, usually in distinct patches

Estimated Volume Recovered: ~0 cm3 (some ~10 µm sized particles). No core recovery

Coring Bit Number: 5

Core Orientation: hade = 3.43°; azimuth = 208.74°; core roll = 289.35°

Sample Serial Numbers: Tube SN233; Seal SN062; Ferrule SNxxx

ACA Temperature at Time of Sealing: 40°C

Estimated Rover-Ambient Pressure and Temperature at Time of Sealing: 749 Pa, 221 K

Estimated Amount of Martian Atmosphere Headspace Gas: 4.9x10<sup>-6</sup> mol

Abrasion Patch Name and Depth: Guillaumes, 8 mm

Anomalous Sample Behavior: Core disintegrated; no recovery

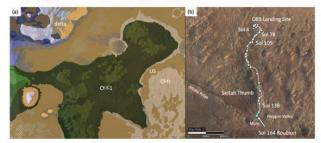
#### **Summary Description**

Collection of Boubling, the first sample target of the Mars 2020 mission, was attempted in the Seitah. Thumb region of the Jezero crater floor (Figure 1). Between Octavia E. Butler landing (OEB) and the sampling site, Perseverance traversed about 1 km southward over nearly continuous low-lying rocks typically forming meter-scale "payerstone" outcrops, with intervening regolith (Figure 2). In HIRISE orbital view, these rocks define a distinctive polygonal pattern, the lower lying of the several expressions of the Crater Floor Fractured Rough (CFF) unit of Stack et al (2020).

Figure 1 | Regional context. (a) geologic map of Sun and Stack (2020) and units of Stack et al. 2020 and (b)

HRISE map with Perseverance traverse path leading to Roughou, Shown for reference are Octavia E. Buttler (OEB)

landing site, Editah Thumb area, Polygon Valley, Mage, and Artub, ridge.



Prior to landing, the CF-fx unit was variously interpreted to be igneous (lava or volcanoclastic) or sedimentary (fluviolacustrine or aeolian) in origin. Although the stratigraphic context of the fractured floor, and indeed the lithology itself, were unknown at the time of sampling, the CF-fx unit was selected for sample acquisition because it is acceptly extensive and because it includes the most heavily cratered terrain to which the rover has access within Jezero crater. As such, a returned sample of this unit was thought to have high science value for understanding the geologic setting and timing of crater floor units, and possibly for calibration of the Mars crater chronology function.

As evidenced by the common whaleback morphology and surface polish and fluting, this expression of Cf-fr. has been eroded to just above ground level by aeolian abrasion (Figure 2). Despite abundant outcrop, little or no visual evidence of sedimentary structure, clasts, or crystals were seen in natural exposures of these rocks. Supercam, data on multiple outcrops along the traverse indicate an altered (hydrated, iron-oxide-bearing) mafic rock with crystal size large enough to create spot-to-spot variability in composition (i.e., > few hundred µm scale). These low-lying Cf-fr rocks are fairly homogenous in appearance and composition along the entire traverse from OEB to Boubjon.

The <u>Roubjan</u> coring attempt, and its companion <u>Guillaumes</u> abraded patch, were undertaken on a low relief rock at the tip of the <u>Séitah</u> Thumb region selected largely to meet first-time sampling

2-2

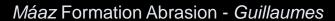
First Science Campaign: Rocks of the Jezero Crater Floor



Proposed previously: fluvio-lacustrine or aeolian sediments, lava flows, impact melt sheet, volcanic ash deposits

(Stack et al., 2020 and references







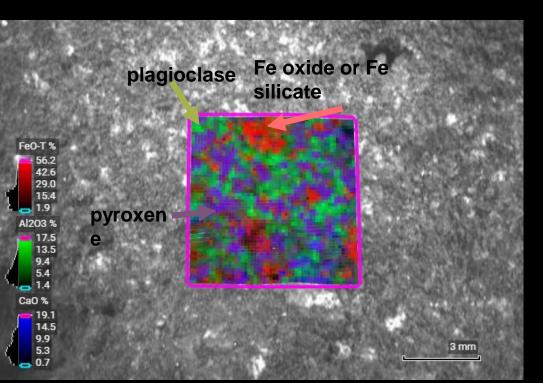


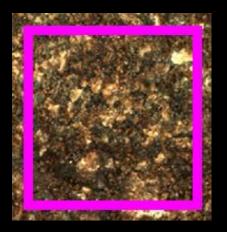
5 mm

**Guillaumes 4 cm merge product** 

Interlocking grain texture, alteration

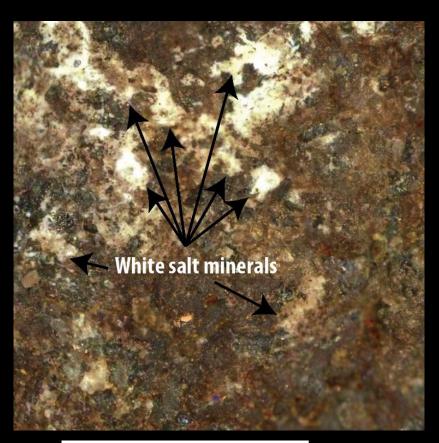
#### Máaz Formation Abrasion - Guillaumes





basaltic mineralogy and bulk composition holocrystalline basalt or microgabbro

## Aqueous alteration in Guillaumes (Máaz Formation)

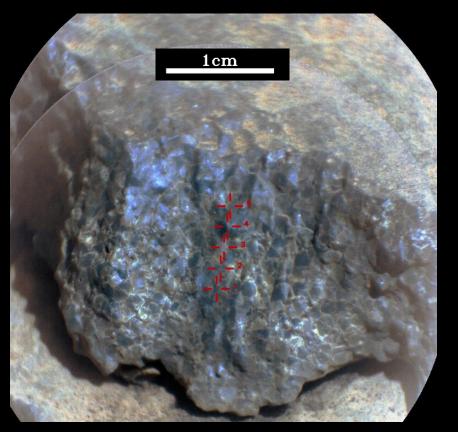


From elemental composition and Raman spectroscopy: Sulfate and perchlorate salts in white patches/vugs

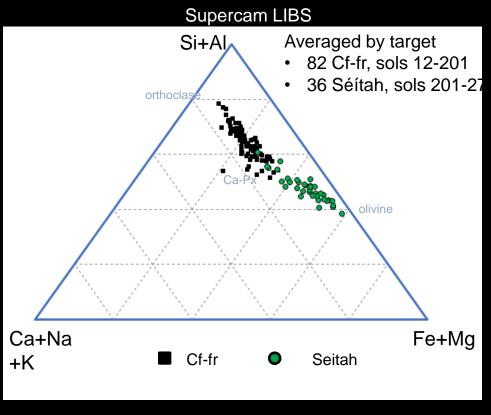
(similar features are seen in the Séítah formation)

4 mm

#### Strong Distinctions Between *Máaz* and *Séítah*



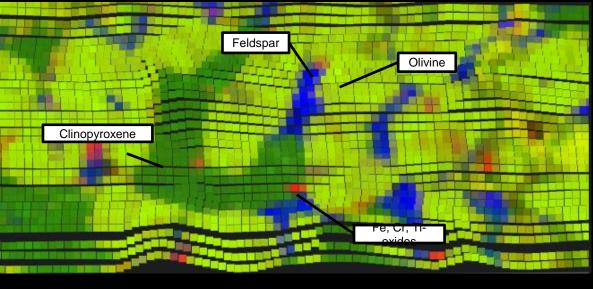
Supercam RMI

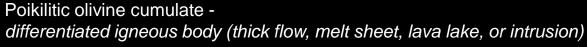


Séítah: abundant olivine and high Mg, with densely packed angular to slightly rounded 2-3 mm-size crystals.



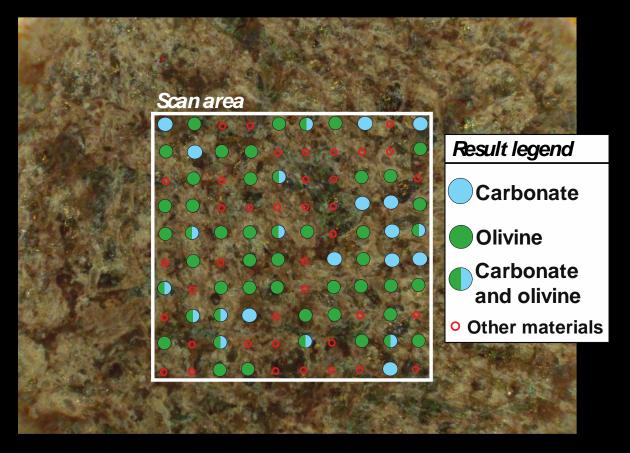
#### Séítah Formation Dourbes Abrasion Patch







#### Raman Detection of Carbonate in Séítah



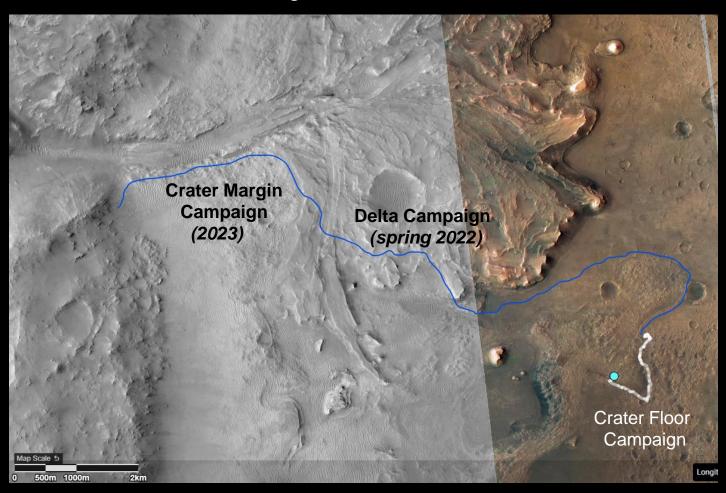
Possible in-situ carbonation of olivine by CO<sub>2</sub> rich water

First Science Campaign: Rocks of the Jezero Crater Floor



Máaz and Séítah are both igneous formations both have experienced multiple styles of aqueous alteration

## Long-term Mars 2020 Plan





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